## Orbitals and Orbital Energies in DFT and (TD)DFT

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## Common misunderstandings

1) The KS orbitals have no physical meaning, they serve only to build the density.

**No:** the orbitals have a better shape and energy than the HF orbitals. They are better suited for qualitative and quantitative MO theory.

2) HF is the best one-electron model (lowest one-det. energy) HF orbitals are the best orbitals / the "real" orbitals
No: HF orbitals and density are too diffuse; *T* (kin. en.) and *V* (el.-nucl. en.) are poor, more in error than *W*

Comment: In HF the one-el. properties are OK, the error is in *W* (el.-el. energy) due to neglect of correlation **No:** errors in one-el. terms *T* and *V* are larger!

3) There is no Koopmans' theorem in DFT. The occupied orbital energies (except HOMO:  $\varepsilon_H = -I_0$ ) are meaningless.

**No:** there is a better-than-Koopmans relation in DFT between (*exact Kohn-Sham*) orbital energies and IPs: deviation for valence of ca. 0.1 eV, against HF ca. 1.1 eV. And theoretically justified!

NB: LDA/GGA orbital energies are very wrong: 4 – 6 eV up

4) The KS HOMO-LUMO gap (in molecules) and the KS band gap (in solids) is wrong (much too small).

No: the KS virtual orbitals have different physical meaning and different orbital energies than HF virtuals.The KS gap (HOMO-LUMO orbital energy difference) is in molecules an excellent approximation for the first excitation energy.

In solids the fundamental gap (I-A) and optical gap (usually close to fundamental gap) are different from the KS band gap for a reason.

(The **reason** is not the derivative discontinuity.)

5) Charge-transfer transitions (excitation out of the HOMO of one molecule to the LUMO of another molecule) are not OK in TDDFT "because of" the derivative discontinuity

**No:** they are more problematic than local excitations because of the physical nature of the KS unoccupied orbitals

6) Computational cost of KS is same as Hartree, much lower than HF.

### **No:** *higher cost* than HF

(unless tricks: density fitting to scale Coulomb part down to  $N^3$  scaling)

Exchange-correlation energy 
$$W_{xc}$$
  

$$E = \langle \Psi | H | \Psi \rangle = T + \int \rho(r) v(r) dr + \frac{1}{2} \int \frac{\rho(r_1) \rho(r_2)}{r_{12}} dr_1 dr_2 + W_{xc}$$

$$W_{XC} = \frac{1}{2} \int dr_1 \rho(r_1) \int \frac{\rho_{XC}^{hole}(r_2 \mid r_1)}{r_{12}} dr_2$$

$$V_{xc}^{hole}(r_1)$$

$$\Gamma(r_1, r_2) = \rho(r_1) \rho(r_2) + \Gamma_{xc}(r_1, r_2) \Rightarrow$$

$$\frac{\Gamma(r_1, r_2)}{\rho(r_1)} = \rho(r_2) + \frac{\Gamma_{xc}(r_1, r_2)}{\rho_{XC}^{hole}(r_2 \mid r_1)}$$

$$(r_1, r_2) = \rho(r_1) + \frac{\Gamma_{xc}(r_1, r_2)}{\rho_{XC}^{hole}(r_2 \mid r_1)}$$

### $W_{xc}$ and Kohn-Sham $E_{xc}$

The Kohn-Sham Molecular Orbital model:

There is a unique local potential  $v_s(\mathbf{r})$  (the KS pot.) such that

$$\hat{h}_{s}\psi_{i}^{s}(\mathbf{r}) = \left\{-\frac{1}{2}\nabla_{\mathbf{r}}^{2} + v_{s}(\mathbf{r})\right\}\psi_{i}^{s}(\mathbf{r}) = \varepsilon_{i}^{s}\psi_{i}^{s}(\mathbf{r})$$

gives KS orbitals such that  $\sum_{i=1}^{H} |\psi_i^s(\mathbf{r})|^2 = \rho^{exact}(\mathbf{r})$ 

$$E = T + \int \rho(\mathbf{r})v(\mathbf{r})d\mathbf{r} + \frac{1}{2}\int \frac{\rho(\mathbf{r}_1)\rho(\mathbf{r}_2)}{r_{12}}d\mathbf{r}_1 d\mathbf{r}_2 + W_{xc}$$
$$E = T_s + \int \rho(\mathbf{r})v(\mathbf{r})d\mathbf{r} + \frac{1}{2}\int \frac{\rho(\mathbf{r}_1)\rho(\mathbf{r}_2)}{r_{12}}d\mathbf{r}_1 d\mathbf{r}_2 + E_{xc}$$

 $\Rightarrow E_{xc} = \underbrace{T - T_s}_{T_c} + W_{xc} \qquad E_{xc}[\rho]: \text{ fentl of } \rho \text{ since } E[\rho] \text{ is.}$ 

$$v_s = v_{nuc} + v_{Coul} + v_{xc}$$

 $v_{xc}(\mathbf{r}) = \frac{\delta E_{xc}}{\delta \rho(\mathbf{r})}$  Not very transparent! Therefore use:

$$E_{xc} = T - T_s + W_{xc} = \int \rho(\mathbf{r}) v_{c,kin}(\mathbf{r}) d\mathbf{r} + \frac{1}{2} \int \rho(\mathbf{r}) v_{xc}^{hole}(\mathbf{r}) d\mathbf{r}$$

$$\frac{\delta E_{xc}}{\delta \rho(\mathbf{r})} = \frac{1}{2} \int \frac{\delta \rho(\mathbf{r}')}{\underbrace{\delta \rho(\mathbf{r})}_{\delta(\mathbf{r}-\mathbf{r}')}} v_{xc}^{hole}(\mathbf{r}') d\mathbf{r}' \qquad \rightarrow v_{xc}^{hole}(\mathbf{r})$$

$$+\int \frac{\delta \rho(\mathbf{r}')}{\delta \rho(\mathbf{r})} v_{c,kin}(\mathbf{r}') d\mathbf{r}' \qquad \rightarrow v_{c,kin}(\mathbf{r})$$

$$+ \int \rho(\mathbf{r}') \frac{\delta}{\delta \rho(\mathbf{r})} \left[ \frac{1}{2} v_{xc}^{hole}(\mathbf{r}') + v_{c,kin}(\mathbf{r}') \right] d\mathbf{r}' \rightarrow v^{resp}(\mathbf{r})$$



# Hartree-Fock errors for bond energies (kcal/mol)

|                       | HF    | Obs.  | Error (% of Obs.) |
|-----------------------|-------|-------|-------------------|
| N <sub>2</sub>        | 115.2 | 228.6 | - 49.6%           |
| $F_2$                 | -37.1 | 38.5  | - 196.4%          |
| H <sub>2</sub> O      | 155.5 | 232.2 | - 33.0%           |
| <b>O</b> <sub>2</sub> | 33.1  | 120.5 | -72.5%            |





Grüning, Gritsenko, Baerends, JCP 118 (2003) 7183

The anomalous F<sub>2</sub> case: RHF energy above two F atoms! Why? Because of long distance RHF error.



Exchange-only (i.e. Hartree-Fock model) yields 98-99% of total energy, so good zero-order model. ("Unfortunately bond energies are of the order of 1-2% so correlation is essential for quantum chemistry")

### A WRONG PERSPECTIVE!

Hartree-Fock (and EXX) relatively OK for some systems: electron gas, atoms -> molecules total energy

Hartree-Fock (and EXX) are lousy zero-order approximation for electron-pair bonding (= chemistry).

LDA and GGA work with localized holes, that is why they easily do better than HF; but not accurate and not along complete dissociation coordinate!

# Hartree-Fock densities are often poor due to bad HF potential: $H_2$

|                                       | $E_{total}^{corr}$ | T <sup>corr</sup> | $V_{el-nuc}^{corr}$ | $W_{el-el}^{corr}$ |
|---------------------------------------|--------------------|-------------------|---------------------|--------------------|
| $H_2 (R = R_e)$                       | -1.1 eV            | +1.3              | -0.5                | -1.9               |
| H <sub>2</sub> ( <i>R</i> =5.0 bohr)  | -3.9               | +8.9              | -8.5                | -4.4               |
| H <sub>2</sub> ( <i>R</i> =10.0 bohr) | -6.3               | +7.9              | -8.6                | -5.6               |

## Hartree-Fock densities are often poor due to bad potential: He, $H_2O$ , Ne, $N_2$

|                  | $E_{total}^{corr}$ | T <sub>kin</sub> | $V_{el-nuc}^{corr}$ | $W_{el-el}^{corr}$ |
|------------------|--------------------|------------------|---------------------|--------------------|
| Не               | -1.1 eV            | +1.1             | -0.1                | -2.1               |
| H <sub>2</sub> O | -7.0               | +6.5             | +1.0                | -14.5              |
| Ne               | -8.9               | +8.3             | +1.4                | -18.5              |
| N <sub>2</sub>   | -11.0              | +13.7            | -13.8               | -11.0              |

## Exact Kohn-Sham orbitals and energies

There is only one KS potential that generates the exact densities (HK): KS orbitals and orbital energies are uniquely defined!

Given exact  $\rho(\mathbf{r})$ , how to find  $v_s$ ?

$$\left\langle \Psi_{2} \left| \hat{H}_{1} \right| \Psi_{2} \right\rangle > \left\langle \Psi_{1} \left| \hat{H}_{1} \right| \Psi_{1} \right\rangle = E_{1}$$
  
$$T_{2} + W_{2} + \int \rho_{2} v_{1} d\mathbf{r} > T_{1} + W_{1} + \int \rho_{1} v_{1} d\mathbf{r}$$
  
and

$$\left\langle \Psi_1 \left| \hat{H}_2 \right| \Psi_1 \right\rangle > \left\langle \Psi_2 \left| \hat{H}_2 \right| \Psi_2 \right\rangle = E_2$$
  
$$T_1 + W_1 + \int \rho_1 v_2 \, d\mathbf{r} > T_2 + W_2 + \int \rho_2 v_2 \, d\mathbf{r}$$

sum up

$$\int \rho_2 v_1 d\mathbf{r} + \int \rho_1 v_2 d\mathbf{r} > \int \rho_1 v_1 d\mathbf{r} + \int \rho_2 v_2 d\mathbf{r}$$

or

$$\int (\rho_2 - \rho_1)(v_2 - v_1) d\mathbf{r} < 0, \text{ i.e. } \int \Delta \rho \Delta v d\mathbf{r} < 0$$

$$\int (\rho_2 - \rho_1)(v_2 - v_1) d\mathbf{r} < 0, \text{ i.e. } \int \Delta \rho \Delta v d\mathbf{r} < 0$$

If in a small region the potential is decreased,  $\Delta v < 0$ , then  $\rho$  must change (cf. HK!),

and  $\Delta \rho$  must be positive over that region, and v.v. (van Leeuwen-Baerends, PRA 1994)

Apply to the KS potential  $v_s$ : by locally adjusting  $v_s$  the density can be made to approach the exact (correlated) density from e.g. CI arbitrarily

closely  $\rightarrow$  generates the *exact* KS potential



Calculate H atom in Gaussian basis: small deviations from exact density.

Generate potential that produces exactly that Gaussian density: small deviations from -1/r, when  $\Delta \rho$  positive  $\Delta v$  negative, and vice versa.

Schipper, Gritsenko, Baerends, Theor. Chem. Acc. **98** (1997) 16

### KS and HF energies of N<sub>2</sub> $D_e=0.37$ a.u.

| R (bohr)                             | 2.074    | 3.0      | 3.5      |
|--------------------------------------|----------|----------|----------|
| $T_s$                                | 109.070  | 108.095  | 108.223  |
| $T_s - T^{HF} =$                     | 0.296    | 0.692    | 0.903    |
| $T - T_s = T_{kin}^{corr}(KS)$       | 0.329    | 0.328    | 0.313    |
| $T - T^{HF} = T^{corr}_{kin}(HF)$    | 0.625    | 1.020    | 1.216    |
|                                      |          |          |          |
| $V_{el-nuc}(\text{exact}=\text{KS})$ | -303.628 | -288.260 | -283.780 |
| $V_{el-nuc}^{corr}(HF)$              | -0.558   | -1.330   | -1.759   |
| W <sub>Coul</sub> (exact)            | 75.068   | 67.858   | 65.666   |
| $W_{Coul}^{corr}(exact - HF)$        | 0.274    | 0.716    | 0.980    |

Gritsenko, Schipper, Baerends, J. Chem. Phys. 107 (1997) 5007

### KS and HF energies of N<sub>2</sub> $D_e=0.37$ a.u.

| R (bohr)                          | 2.074   | 3.0     | 3.5     |
|-----------------------------------|---------|---------|---------|
| $W_X$ (KS orbitals)               | -13.114 | -12.621 | -12.490 |
| $W_X - W_X^{HF} =$                | 0.006   | -0.040  | -0.067  |
| $W_c = W_{XC} - W_X$              | -0.804  | -0.969  | -1.063  |
| $W_{c}(HF) = W_{XC} - W_{X}^{HF}$ | -0.810  | -1.009  | -1.124  |
|                                   |         |         |         |
| $E_c(KS)$                         | -0.475  | -0.641  | -0.750  |
| $E_c(\mathrm{HF})$                | -0.469  | -0.603  | -0.687  |
| $E_c(KS) - E_c(HF)$               | -0.006  | -0.0038 | -0.063  |

Gritsenko, Schipper, Baerends, J. Chem. Phys. 107 (1997) 5007

### Definition of correlation energy



Conclusion HF versus KS det.

 $-\Psi^{\rm HF} \text{ better total energy (marginally)}$  $E_c \le E_c^{\ HF}$ 

 $\Psi_{s}$  better for:

 $\begin{cases} V_{el-nuc} \\ W_{Coul} \end{cases}$  no correlation error  $T_s$ : (much) smaller correl. error

HF "distorts" density (more diffuse) if: gain by lowering  $T^{HF}$  is larger (even if barely) than loss by less stable V

### Hartree-Fock: good for atoms, not for molecules (bonds)

In an electron pair bond:

- a) HF orbitals will be too diffuse (density too diffuse)
- $\rightarrow$  kinetic energy too low
- $\rightarrow$  electron-nuclear energy too high (not negative enough)

b) this is worse in case of multiple bonds

c) common statement

"one-particle properties (also the electron density!) are good in the Hartree-Fock model, it is the el.-el. interaction that is wrong, because of lack of electron correlation (electrons do not avoid each other sufficiently, cf. the presence of ionic configurations in the  $H_2$  wavefunction)"

### **IS WRONG**

# What is the meaning of KS orbital energies?

Prevailing view, see e.g.

R.G. Parr, W. Yang, *DFT of Atoms and Molecules*, 1989:

"..one should expect no simple physical meaning for the KS orbital energies. *There is none*."

### KS and HF orbital energies and VIPs for H $_2O$

| H <sub>2</sub> O | MO     | HF               | KS               | Expt.   | $I_k + \varepsilon_k$ | $\sqrt{\varepsilon_N - \varepsilon_k}$ |
|------------------|--------|------------------|------------------|---------|-----------------------|--|
|                  |        | $-\varepsilon_k$ | $-\varepsilon_k$ | $I_k$   |                       |  |
|                  | $1b_1$ | 13.76            | 12.63            | 12.62   | -0.01                 |  |
|                  | $3a_1$ | 15.77            | 14.78            | 14.74   | -0.04                 |  |
|                  | $1b_2$ | 19.29            | 18.46            | 18.55   | 0.09                  |  |
| Average          |        | 0.97             | 0.05             |         |                       |  |
| Dev.             |        |                  |                  |         |                       |  |
|                  |        |                  |                  |         |                       |  |
|                  | $2a_1$ | 36.48            | 30.89            | 32.2    | 1.31                  | 4.27                                   |
|                  | $1a_1$ | 559.37           | 516.96           | 539.90  | 22.94                 | 22.46                                  |
| Average          |        | 11.88            | 12.13            | Chong,  | Braïda,Gr             | itsenko,                               |
| Dev.             |        |                  |                  | Baerenc | ls, JCP 20            | 02-2004                                |

### HCl: KS, BP and HF orbital energies and VIPs

| HC1      | M         | HF               | GGA-BP             | KS               | Expt. |
|----------|-----------|------------------|--------------------|------------------|-------|
|          | Ο         | $-\mathcal{E}_i$ | $-\mathcal{E}_i$   | $-\mathcal{E}_i$ | $I_i$ |
|          | $2\pi$    | 12.97            | 8.13 (+4.64=12.77) | 12.77            | 12.77 |
|          | $5\sigma$ | 17.04            | 11.90 (16.53)      | 16.53            | 16.6  |
|          | $4\sigma$ | 30.41            | 21.22 (25.86)      | 25.82            | 25.8  |
| AAD(val) |           | 1.75             | 4.68 (0.04)        | 0.03             |       |
|          | $1\pi$    | 218.77           | 190.98 (195.62)    | 199.59           |       |
|          | 3σ        | 218.84           | 191.27 (195.91)    | 199.79           |       |
|          | $2\sigma$ | 287.75           | 250.44 (255.08)    | 259.80           |       |

## $v_{xc}^{hole}$ in KS pot.: consequences

1) Good shape of occupied orbitals and density, good orbital energies

2) Good shapes and energies of virtual orbitals. Consequences for excitation energies etc. (TDDFT): excitations from occupied orbitals  $\varphi_i$  (*i*,*j*,*k*,... for occ. orbitals) to unoccupied orbitals to  $\varphi_a$  (*a*,*b*,*c*,*d*.. for unocc. (virtual) orbitals)

What are virtuals like in DFT? And in Hartree-Fock?

Big difference between HF and KS virtuals: necessary to understand the difference to understand

- why TDDFT works so well (in general for molecules);
- why there is a problem with charge-transfer transitions
- the "bandgap problem" in solids

### HF, DFA and exact KS HOMO orbital energies

|                  | HF     | LDA    | BLYP   | KS = -I |
|------------------|--------|--------|--------|---------|
| H <sub>2</sub>   | -16.18 | -10.26 | -10.39 | -16.44  |
| H <sub>2</sub> O | -13.88 | -7.40  | -7.21  | -12.62  |
| HF               | -17.69 | -9.82  | -9.64  | -16.19  |
| N <sub>2</sub>   | -16.71 | -11.89 | -11.49 | -16.68  |
| СО               | -15.1  | -9.11  | -9.00  | -14.01  |
| HCN              | -13.50 | -9.23  | -8.87  | -13.61  |
| FCN              | -13.65 | -8.97  | -8.62  | -13.67  |
| HC1              | -12.98 | -8.15  | -7.91  | -12.77  |
|                  |        |        |        |         |

#### KS HOMO is equal to – I;

HF HOMO is appr. equal to – I (frozen orbital approx.) LDA, GGA orbital energies are upshifted by ca. 4.5 eV (uniformly: occup. and unoccup *valence* orbitals)

### HF, DFA and exact KS LUMO orbital energies

|                  | HF    | LDA   | BLYP  | KS    |
|------------------|-------|-------|-------|-------|
| H <sub>2</sub>   | +1.42 | +0.31 | +0.12 | -3.93 |
| H <sub>2</sub> O | +0.80 | -0.92 | -1.06 | -5.11 |
| HF               | +0.81 | -0.93 | -1.13 | -5.71 |
| N <sub>2</sub>   | +3.91 | -2.21 | -1.91 | -6.77 |
| СО               | +1.88 | -2.24 | -1.94 | -6.56 |
| HCN              | +1.93 | -1.33 | -1.07 | -5.53 |
| FCN              | +1.16 | -1.66 | -1.59 | -6.01 |
| HC1              | +0.79 | -1.11 | -1.15 | -5.36 |
|                  |       |       |       |       |

KS LUMO is at negative energy: a bound one-electron state in the KS potential.

HF LUMO is most of the time *unbound* (positive orbital energy) LDA,GGA LUMO: still negative -> therefore bound state Meaning of unoccupied orbital energies  $\varepsilon_a$ ,  $\varepsilon_b$ , ....

*HF*: unocc. orbital represents *added* electron  $\rightarrow \varepsilon_a^{HF}$  is *affinity* level;  $\varepsilon_a^{HF} - \varepsilon_i^{HF}$  is NOT excitation energy

*KS*: unocc. orbital represents *excited* electron  $\rightarrow \varepsilon_a^{KS} - \varepsilon_i^{KS}$  IS good appr. to excitation energy



HF virtual orbitals are at (much) higher energy and (way) more diffuse than KS virtual orbitals

## KS HOMO-LUMO gaps $\Delta$ are excellent approx. to excitation energies

|                  | $\Delta^{\rm HF}$ | Δ <sup>LDA</sup> | $\Delta^{\text{BLYP}}$ | Δ <sup>KS</sup> | Expt. exc<br>singlet | it. energy<br>triplet |
|------------------|-------------------|------------------|------------------------|-----------------|----------------------|-----------------------|
| H <sub>2</sub>   | 17.6              | 10.6             | 10.5                   | 12.5            | 12.7                 | 11.7                  |
| H <sub>2</sub> O | 14.7              | 6.5              | 6.2                    | 7.5             | 7.65                 | 7.5                   |
| HF               | 18.5              | 8.9              | 8.5                    | 10.5            | 10.3                 | 9.9                   |
| N <sub>2</sub>   | 19.9              | 9.7              | 9.6                    | 9.9             | 9.3-10.3             | 7.8-8.9               |
| СО               | 17.0              | 6.9              | 7.1                    | 7.5             | 8.5                  | 6.3                   |
| HCN              | 15.4              | 7.9              | 7.8                    | 8.0             | 8.8                  | 6.2                   |
| FCN              | 14.8              | 7.3              | 7.0                    | 7.6             | 8.4                  | 7.8                   |
| HCl              | 13.8              | 7.0              | 6.8                    | 7.4             | 7.8                  | 7.4                   |
|                  |                   |                  |                        |                 |                      |                       |

The LDA, GGA gaps are similar (slightly smaller) than KS gaps
 the upshift is similar for HOMO and (a bit smaller for) LUMO
 HF gaps are much larger: they are Koopmans' approx. IP – EA



# What is the meaning of a LUMO with positive orbital energy?





# What is the meaning of HF LUMO with positive energy?

Note: positive one-electron states in a potential (zero at infinity):

- there is a continuum of positive states;
- most have plane-wave behavior with only a few orthogonality wiggles over the molecular region;
- at specific energy (small energy ranges) the one-electron states have large amplitude in the molecular region (small plane-wave like outside)
  -> "scattering resonances" with resonance energies corresponding to potential electron capture to form a temporary negative ion, which will decay after some time to molecule + free electron.

Since energy at scattering resonance is *positive*,

i.e. *higher* than free molecule and electron: *negative electron affinity*!

If there are no negative energy unoccupied orbitals (bound states) for the HF operator (frequently!), what is the meaning of the pos. energy orbitals?



Calculated LUMO energies vs EA for 113 molecules. EA from CCSD(T), basis: 6-311++G(3df,3dp)

(Kar, Song, Hirao, JCC 2013) Almost all HF  $\mathcal{E}_{LUMO}$  positive!

## Orbital energies (eV) of the positive energy HF LUMO of $H_2$ as function of the basis (STOs)

|                   | SZ     | DZ     | DZP    | TZP    | TZ2P   | QZ4P   | ETQZ3P<br>2D |
|-------------------|--------|--------|--------|--------|--------|--------|--------------|
| $1\sigma_{\rm u}$ | 18.12  | 5.52   | 5.11   | 3.39   | 3.45   | 2.67   | 1.18         |
| $1\sigma_{g}$     | -15.88 | -16.26 | -16.20 | -16.21 | -16.20 | -16.18 | -16.18       |
| gap               | 34.00  | 21.77  | 21.31  | 19.59  | 19.66  | 18.85  | 17.36        |
|                   |        |        |        |        |        |        |              |

## Orbital energies (eV) of the positive energy HF LUMO of $H_2$ as function of the basis (Gaussians)

|                   | cc-<br>pVDZ | сс-<br>pVQZ | cc-<br>pV5Z | aug-cc-<br>pVDZ | aug-cc-<br>pVTZ | aug-cc-<br>pVQZ | aug-cc-<br>pV5Z |
|-------------------|-------------|-------------|-------------|-----------------|-----------------|-----------------|-----------------|
| $1\sigma_{\rm u}$ | 5.372       | 3.91        | 3.14        | 1.67            | 1.42            | 1.28            | 1.14            |
| $1\sigma_{g}$     | -16.11      | -16.18      | -16.18      | -16.12          | -16.18          | -16.18          | -16.18          |
| gap               | 21.48       | 20.66       | 19.33       | 17.80           | 17.60           | 17.46           | 17.32           |

Orbital energies of LUMO are arbitrary; completely determined by the basis set.

Go to zero for complete basis.

What about shape? Should go to infinitely extended.

# Shape of the $1\sigma_u$ LUMO density of H<sub>2</sub> as a function of basis set:



# Practical ways to get scattering resonances (negative EAs) with basis set calculations

Stabilization method (H. S. Taylor et al.), also called SKT (stabilization Koopmans' method):

Systematically scan through the spectrum of positive energies by scaling the coefficients of all diffuse basis functions to very low value (diffuse). Then orbital energies go down in energy as function of scaling parameter  $\alpha$ .

Detect resonance energies by inspecting the orbitals; when getting high amplitude in molecular region, you are at resonance energy.

Or by looking at curves of orbital energy as function of  $\alpha$ : resonance energies show up as "avoided crossing".

See K. Jordan et al. (JPC-A **104** (2000) 9605) and Cheng et al. (JPC-A **116** (2012) 12364)

Orbital energies and excitation energy calculations (TDDFT)

TDDFT: 
$$\left(\boldsymbol{\mathcal{E}}^{2} + 2\sqrt{\boldsymbol{\mathcal{E}}}K\sqrt{\boldsymbol{\mathcal{E}}}\right)\mathbf{F}_{q} = \omega_{q}^{2}\mathbf{F}_{q}$$
  
 $\left(\boldsymbol{\mathcal{E}}^{2}\right)_{ia,jb} = \delta_{ij}\delta_{ab}(\varepsilon_{a} - \varepsilon_{i})^{2}$ 

*K* is "coupling matrix", see later

Suppose 
$$i \to a$$
 does not couple to other  $j \to b$   
(single pole approximation, SPA),  $q \cong i \to a$   
 $\left[ (\varepsilon_a - \varepsilon_i)^2 + 2(\varepsilon_a - \varepsilon_i) \int \varphi_i(\mathbf{r}) \varphi_a(\mathbf{r}) f_{xc}(\mathbf{r}, \mathbf{r}') \varphi_i(\mathbf{r}') \varphi_a(\mathbf{r}') d\mathbf{r} d\mathbf{r}' \right] \mathbf{F}_q = \omega^2 \mathbf{F}_q$   
 $\Rightarrow \omega \approx (\varepsilon_a - \varepsilon_i) + \underbrace{\langle \varphi_i \varphi_a | f_{xc} | \varphi_i \varphi_a \rangle}_{small}$ 

 $(\varepsilon_a - \varepsilon_i) \approx$  excitation energy (in molecules!)

## KS HOMO-LUMO gaps $\Delta$ are excellent approx. to excitation energies

|                  | $\Delta^{\rm HF}$ | Δ <sup>LDA</sup> | $\Delta^{\text{BLYP}}$ | Δ <sup>KS</sup> | Expt. exc<br>singlet | it. energy<br>triplet |
|------------------|-------------------|------------------|------------------------|-----------------|----------------------|-----------------------|
| H <sub>2</sub>   | 17.6              | 10.6             | 10.5                   | 12.5            | 12.7                 | 11.7                  |
| H <sub>2</sub> O | 14.7              | 6.5              | 6.2                    | 7.5             | 7.65                 | 7.5                   |
| HF               | 18.5              | 8.9              | 8.5                    | 10.5            | 10.3                 | 9.9                   |
| N <sub>2</sub>   | 19.9              | 9.7              | 9.6                    | 9.9             | 9.3-10.3             | 7.8-8.9               |
| СО               | 17.0              | 6.9              | 7.1                    | 7.5             | 8.5                  | 6.3                   |
| HCN              | 15.4              | 7.9              | 7.8                    | 8.0             | 8.8                  | 6.2                   |
| FCN              | 14.8              | 7.3              | 7.0                    | 7.6             | 8.4                  | 7.8                   |
| HCl              | 13.8              | 7.0              | 6.8                    | 7.4             | 7.8                  | 7.4                   |
|                  |                   |                  |                        |                 |                      |                       |

The LDA, GGA gaps are similar (slightly smaller) than KS gaps
 the upshift is similar for HOMO and (a bit smaller for) LUMO
 HF gaps are much larger: they are Koopmans' approx. IP – EA



proper KS pot. important for virtual orbital spectrum!



#### Acetone: orbital energy differences and excitation energies (eV)

| Funct.         | State           | Weight | Ei     | Ea    | $\Delta \varepsilon_{ia}$ | ω    | $\omega - \Delta \varepsilon_{ia}$ | $\omega - E_{exp}$ |
|----------------|-----------------|--------|--------|-------|---------------------------|------|------------------------------------|--------------------|
| SAOP           | $1A_2$          | 1.00   | -10.25 | -5.92 | 4.33                      | 4.59 | 0.26                               | 0.16               |
| ("exact"       | 1B <sub>2</sub> | 1.00   | -10.25 | -4.18 | 6.07                      | 6.09 | 0.02                               | -0.27              |
| KS)            | $2A_2$          | 0.84   | -10.25 | -2.72 | 7.53                      | 7.52 | 0.00                               | 0.16               |
|                | 2A <sub>1</sub> | 0.97   | -10.25 | -3.09 | 7.16                      | 7.21 | 0.05                               | -0.20              |
|                | 2B <sub>2</sub> | 0.97   | -10.25 | -2.63 | 7.62                      | 7.64 | 0.02                               | 0.15               |
|                | 3A <sub>1</sub> | 0.97   | -10.25 | -2.04 | 8.21                      | 8.20 | 0.00                               | 0.40               |
|                | 3B <sub>2</sub> | 0.97   | -10.25 | -2.51 | 7.74                      | 7.74 | 0.00                               | -0.35              |
|                | 1B <sub>1</sub> | 0.95   | -10.25 | -1.83 | 8.42                      | 8.43 | 0.01                               | 0.26               |
| BP86           | $1A_2$          | 1.00   | -5.71  | -1.70 | 4.01                      | 4.27 | 0.26                               | -0.16              |
|                | 1B <sub>2</sub> | 1.00   | -5.71  | -0.61 | 5.10                      | 5.10 | 0.00                               | -1.26              |
|                | $2A_2$          | 1.00   | -5.71  | -0.11 | 5.60                      | 5.59 | 0.00                               | -1.77              |
|                | 2A <sub>1</sub> | 1.00   | -5.71  | -0.13 | 5.58                      | 5.58 | 0.00                               | -1.83              |
| <b>JCTC 10</b> | 2B <sub>2</sub> | 1.00   | -5.71  | -0.07 | 5.64                      | 5.64 | - 0.01                             | -1.85              |
| 4432           | 3A <sub>1</sub> | 0.98   | -5.71  | +0.36 | 6.07                      | 6.06 | - 0.01                             | -1.74              |
| (2014)         | 3B <sub>2</sub> | 1.00   | -5.71  | +0.05 | 5.76                      | 5.75 | 0.00                               | -2.34              |
| L              | 1B <sub>1</sub> | 1.00   | -5.71  | +0.31 | 6.02                      | 6.01 | -0.01                              | -2.16              |

| Funct.         | State           | Weight | $\mathcal{E}_i$ | $\mathcal{E}_a$ | $\Delta \varepsilon_{ia}$ | ω    | $\omega - \Delta \varepsilon_{ia}$ | $\omega - E_{exp}$ |
|----------------|-----------------|--------|-----------------|-----------------|---------------------------|------|------------------------------------|--------------------|
| HF             | $1A_2$          | 0.47   | -11.23          | +3.96           | 15.18                     | 5.03 | -10.15                             | 0.60               |
|                | 1B <sub>2</sub> | 0.36   | -11.23          | +0.62           | 11.85                     | 8.24 | -3.61                              | 1.88               |
|                | $2A_2$          | 0.43   | -11.23          | +1.02           | 12.25                     | 9.02 | -3.23                              | 1.66               |
|                | 2A <sub>1</sub> | 0.20   | -11.23          | +0.96           | 12.19                     | 9.07 | -3.12                              | 1.66               |
|                | 2B <sub>2</sub> | 0.31   | -11.23          | +1.20           | 12.43                     | 9.13 | -3.30                              | 1.64               |
|                | 3A <sub>1</sub> | 0.21   | -13.20          | +3.96           | 17.15                     | 9.41 | -7.74                              | 1.61               |
|                | 3B <sub>2</sub> | 0.23   | -11.23          | +1.74           | 12.96                     | 9.59 | -3.37                              | 1.50               |
|                | 1B <sub>1</sub> | 0.29   | -15.23          | +3.96           | 19.18                     | 9.66 | -9.52                              | 1.49               |
| M06-2X         | $1A_2$          | 0.52   | -8.85           | +0.78           | 9.63                      | 4.03 | -5.60                              | -0.40              |
|                | $1B_2$          | 0.73   | -8.85           | -0.34           | 8.51                      | 6.54 | -1.97                              | 0.18               |
|                | 2A <sub>2</sub> | 0.62   | -8.85           | +0.04           | 8.88                      | 7.33 | -1.55                              | -0.03              |
|                | 2A <sub>1</sub> | 0.62   | -8.85           | +0.03           | 8.87                      | 7.38 | -1.49                              | -0.03              |
| <b>JCTC 10</b> | 2B <sub>2</sub> | 0.45   | -8.85           | +0.15           | 9.00                      | 7.40 | - 1.60                             | -0.09              |
| 4432           | 3A <sub>1</sub> | 0.79   | -8.85           | +0.74           | 9.58                      | 8.03 | - 1.55                             | 0.23               |
| (2014)         | 3B <sub>2</sub> | 0.42   | -8.85           | +0.64           | 9.49                      | 7.80 | -1.69                              | -0.29              |
|                | $1B_1$          | 0.92   | -8.85           | +0.73           | 9.58                      | 8.12 | -1.45                              | -0.05              |

#### Acetone: orbital energy differences and excitation energies (eV)

occupied orbital energies LDA/GGA: typically 4 – 6 eV too high (HF typically 1 eV too low) (an error of 100 – 140 kcal/mol, totally unacceptable in the total energy!)

Why?

- Not because of wrong asymptotics of potentials
- Not because of self-interaction error

But because LDA/GGA potentials are strongly upshifted in the bulk molecular region

### Exact and LDA KS potentials



Almbladh and Pedroza, PRB 1985

### Exact and LDA KS potentials



van Leeuwen and Baerends, PRA 1994



Grüning, Gritsenko, Baerends, JCP 2002

Electron Gas (EG) exchange energy  

$$E_X^{LDA} = \int \rho(\mathbf{r}) \varepsilon_X^{LDA}(\mathbf{r}) d\mathbf{r} = \frac{1}{2} \int \rho(\mathbf{r}) v_{Xhole}^{EG}(\mathbf{r}) d\mathbf{r}$$

$$v_{Xhole}^{EG}(\mathbf{r}) = -3\left(\frac{3}{8\pi}\right)^{\frac{1}{3}}\rho(\mathbf{r})^{\frac{1}{3}}$$

Slater: square hole around position **r** of electron with depth  $-(1/2)\rho(\mathbf{r})$ , integrating to -1, gives practically same potential  $C\rho(\mathbf{r})^{1/3}$ 

Slater (band structure theorists) applied this in one-electron equations as exchange potential

### Gaspar-Kohn-Sham (DFT):

Proper variational derivation of one-el. equations for the optimal orbitals yields: DA = 2 EG

$$v_X^{LDA}(\mathbf{r}) = \frac{2}{3} v_{Xhole}^{EG}(\mathbf{r})$$

 $\rightarrow$ Slater's X $\alpha$  method: use potential  $\alpha . v_{Xhole}^{EG}(\mathbf{r})$ 

 $\alpha$  has been determined in various ways in atoms (e.g. K. Schwarz, 1972):

- exact exchange energy optimized;
- Virial Theorem obeyed
- energy equal to Hartree-Fock energy

 $\rightarrow \alpha$  in range 0.78 (lightest elements) -0.71 (beyond first row)

or in molecules (E. J. Baerends, 1973) ( $\alpha = 0.70$ )

### $\rightarrow$ Slater's X $\alpha$ method: use potential $\alpha v_{Xhole}^{EG}$

Actually, Slater resisted initially: - orbital energies are "better" with full Slater (EG) exchange pot.!

J.C. Slater, *Quantum Theory of Molecules and Solids*, Vol. 3 (1967): *Insulators, semiconductors and metals*p. 243: Some writers have felt that on account of the importance of the variation principle for the toal energy, it was more important to use ... an exchange correction determined by variation of the total energy ...

The author does not agree with this point of view, feeling that the oneelectron energies are more important in energy-band calculations.

The slightly different wavefunctions which we find by use of [the full  $v_{Xhole}^{EG}(\mathbf{r})$ ] rather than  $(2/3)v_{Xhole}^{EG}(\mathbf{r})$  will only make a second order change in the total energy.

### Orbital energies from Slater (EG), LDA, GGA compared to exact KS

| N <sub>2</sub>               | $3\sigma_{g}$ | 1π <sub>u</sub> | $2\sigma_{u}$ |
|------------------------------|---------------|-----------------|---------------|
| $-I_i$                       | - 15.58       | - 16.83         | - 18.75       |
| $\mathcal{E}_i$ (KS)         | - 15.58       | - 16.84         | - 18.89       |
| $\mathcal{E}_i$ (Slater)     | - 15.39       | - 16.66         | - 18.62       |
| $\mathcal{E}_i(\text{BLYP})$ | - 10.28       | - 11.49         | - 13.39       |
|                              |               |                 |               |

#### Orbital energies from Slater (EG), LDA, GGA compared to exact KS

| H <sub>2</sub> CO              | 2b <sub>2</sub> | 1b <sub>1</sub> | 5a <sub>1</sub> | 1b <sub>2</sub> | <b>4</b> a <sub>1</sub> |
|--------------------------------|-----------------|-----------------|-----------------|-----------------|-------------------------|
| $-I_i$                         | - 10.9          | - 14.5          | - 16.1          | - 17.0          | -21.4                   |
| $\mathcal{E}_{i}(\mathrm{KS})$ | - 10.90         | - 14.26         | - 15.51         | - 16.46         | - 20.04                 |
| $\mathcal{E}_i$ (Slater)       | - 10.99         | - 14.56         | - 15.91         | - 16.82         | - 20.57                 |
| $\mathcal{E}_i(\text{BLYP})$   | - 6.22          | - 9.94          | - 10.93         | - 12.08         | - 15.59                 |
|                                |                 |                 |                 |                 |                         |

## The LDA X-response pot. is (way) too repulsive $E_X^{LDA} = \int \rho(\mathbf{r}) \varepsilon_X^{LDA}(\mathbf{r}) d\mathbf{r} = \frac{1}{2} \int \rho(\mathbf{r}) v_{Xhole}^{EG}(\mathbf{r}) d\mathbf{r}$

$$v_X^{LDA}(\mathbf{r}) = \frac{\delta E_X^{LDA}}{\delta \rho(\mathbf{r})} = v_{Xhole}^{EG}(\mathbf{r}) + v_{Xresp}^{LDA}(\mathbf{r})$$

$$v_X^{LDA}(\mathbf{r}) = -2\left(\frac{3}{8\pi}\right)^{\frac{1}{3}} \rho(\mathbf{r})^{\frac{1}{3}} = -3\left(\frac{3}{8\pi}\right)^{\frac{1}{3}} \rho(\mathbf{r})^{\frac{1}{3}} + \left(\frac{3}{8\pi}\right)^{\frac{1}{3}} \rho(\mathbf{r})^{\frac{1}{3}}$$
$$v_{Xhole}^{EG}(\mathbf{r}) + v_{Xresp}^{LDA}(\mathbf{r})$$
$$(1/3) \text{ of hole pot.!}$$
$$(1/2) \text{ of total X pot.}$$

## What is the right response potential?

$$v_s = v_{nuc} + v_{Coul} + v_{xc}(\mathbf{r})$$

$$v_{xc}(\mathbf{r}) = v_{xc}^{hole}(\mathbf{r}) + v_{c,kin}(\mathbf{r}) + v^{resp}(\mathbf{r})$$

$$\downarrow \qquad \qquad \downarrow \qquad \qquad \downarrow$$
energy density for  $W_{xc}$  for  $T_c$ 

## What is the right response potential?

$$v_s = v_{nuc} + v_{Coul} + v_{xc}(\mathbf{r})$$

$$v_{xc}(\mathbf{r}) = v_{xc}^{hole}(\mathbf{r}) + v_{c,kin}(\mathbf{r}) + v^{resp}(\mathbf{r})$$
  

$$\downarrow \qquad \qquad \downarrow \qquad \qquad \downarrow$$
  
energy density for  $W_{xc}$  for  $T_c$ 

$$v^{resp}(\mathbf{r}) = \int \rho(\mathbf{r}') \frac{\delta v_{c,kin}(\mathbf{r}')}{\delta \rho(\mathbf{r})} d\mathbf{r}' + \frac{1}{2} \int \rho(\mathbf{r}') \rho(\mathbf{r}'') \frac{\delta g(\mathbf{r}',\mathbf{r}'')}{\delta \rho(\mathbf{r})} d\mathbf{r}' d\mathbf{r}''$$

exact expression for  $v^{resp}$ :  $v^{N-1} - v_s^{N-1}$ 

$$v^{resp}(1) = v^{N-1}(1) - v_s^{N-1}(1) = \sum_{j=1}^{N-1} \frac{\left| d_j(1) \right|^2}{\rho(1)} \left( I_j - I_0 \right) - \sum_{j=1}^{N-1} \frac{\left| \psi_j^s(1) \right|^2}{\rho(1)} \left( -\varepsilon_j - I_0 \right)$$

Dyson orbitals: very similar to KS orbitals and HF orbitals for primary ion states exact ionization energies

The response potential consists of contributions that have step like behavior when going from one shell to the next in an atom or molecule

## What is the right response potential?

Krieger-Li-Iafrate (1994) derived a good approximation to the OEP potential in the exact-exchange only case:

$$v_X^{KLI}(\mathbf{r}) \approx v_{Xav}^{HF} + \sum_{i=1}^H w_i \frac{|\psi_i(\mathbf{r})|^2}{\rho(\mathbf{r})}$$

$$v_{Xav}^{HF}(\mathbf{r}) = \sum_{i}^{H} \frac{|\psi_i(\mathbf{r})|^2 v_{Xi}^{HF}(\mathbf{r})}{\rho(\mathbf{r})} \quad \text{a.k.a.} \quad v_X^{Slater}(\mathbf{r})$$

Note 
$$E_X^{HF} = \frac{1}{2} \int \rho(\mathbf{r}) v_{Xav}^{HF}(\mathbf{r}) d\mathbf{r}$$

## Approximation (GLLB) of $v_{Xresp}$

Step behavior introduced with  $w_i = f(\varepsilon_F - \varepsilon_i)$ 

scaling density  $\rho_{\lambda}(\mathbf{r}) = \lambda^{3} \rho(\lambda \mathbf{r})$  yields  $v_{\chi}[\rho_{\lambda}](\mathbf{r}) = \lambda v_{\chi}[\rho](\lambda \mathbf{r})$ 

dictates, with  $\varepsilon_i[\rho_{\lambda}] = \lambda^2 \varepsilon_i[\rho]$ ,

$$f(\lambda^{2}(\varepsilon_{F} - \varepsilon_{i})) = \lambda f(\varepsilon_{F} - \varepsilon_{i}) \text{ so } f \text{ must be prop. to square root:}$$
$$f \rightarrow K \sqrt{\varepsilon_{F} - \varepsilon_{i}}$$
$$v_{Xresp}^{\text{model}}(\mathbf{r}) = K[\rho] \sum_{i=1}^{H} \sqrt{\varepsilon_{F} - \varepsilon_{i}} \frac{|\psi_{i}(\mathbf{r})|^{2}}{\rho(\mathbf{r})}$$

Exact in EG with K = 0.382



### Response potential in Be

R. van Leeuwen, O. Gritsenko, E. J. Baerends, Z. Phys. D 33 (1995) 229

## Steps in the (exchange) pot. of Krypton come from response part of pot.



O. Gritsenko, R. van Leeuwen, E. J. Baerends J. Chem. Phys. **101** (1994) 8955

## N<sub>2</sub> orbital energies

| N <sub>2</sub> (K=0.382)           | $2\sigma_{\mathbf{g}}$ | $2\sigma_{\mathbf{u}}$ | 2π <sub>u</sub> | 3σ <sub>g</sub> (HOMO) |
|------------------------------------|------------------------|------------------------|-----------------|------------------------|
| <i>v<sub>xhole,ii</sub></i> (LDA)  | - 29.47                | -24.72                 | - 23.51         | - 24.89                |
| $v_{xhole,ii}$ (B)                 | - 1.83                 | -2.52                  | -2.48           | -2.54                  |
| <i>v<sub>xresp,ii</sub></i> (GLLB) | + 6.05                 | + 3.97                 | + 5.10          | + 4.56                 |
| $v_{c,ii}$ (VWN)                   | - 1.91                 | - 1.76                 | - 1.75          | - 1.76                 |
| <sup>V</sup> xctot,ii              | - 27.16                | - 25.03                | - 22.63         | - 24.63                |
| $h_{ii} + v_{H,ii}$                | - 6.17                 | + 5.73                 | + 6.16          | + 8.89                 |
| $\boldsymbol{\mathcal{E}}_i$       | - 33.33                | - 19.30                | - 16.47         | - 15.74                |
| <b>e</b> <sub>i</sub> (KS)         | - 33.67                | - 18.89                | - 16.84         | - 15.58                |
| $\Delta \varepsilon_i$             | 0.34                   | -0.41                  | 0.37            | -0.16                  |

## N<sub>2</sub> orbital energies

| N <sub>2</sub> (K=0.382)           | $2\sigma_{g}$ | $2\sigma_{\mathbf{u}}$ | $2\pi_{\mathbf{u}}$ | 3σ <sub>g</sub> (HOMO) |
|------------------------------------|---------------|------------------------|---------------------|------------------------|
| v <sub>xhole,ii</sub> (LDA)        | - 29.47       | -24.72                 | - 23.51             | - 24.89                |
| $v_{xhole,ii}$ (B)                 | - 1.83        | -2.52                  | -2.48               | -2.54                  |
| <i>v<sub>xresp,ii</sub></i> (GLLB) | + 6.05        | + 3.97                 | + 5.10              | + 4.56                 |
| $v_{c,ii}$ (VWN)                   | - 1.91        | - 1.76                 | - 1.75              | - 1.76                 |
| <sup>V</sup> xctot,ii              | - 27.16       | - 25.03                | - 22.63             | - 24.63                |
| $h_{ii} + v_{H,ii}$                | - 6.17        | + 5.73                 | + 6.16              | + 8.89                 |
| <i>E</i> <sub>i</sub>              | - 33.33       | - 19.30                | - 16.47             | - 15.74                |
| <b>E</b> <sub>i</sub> (KS)         | - 33.67       | - 18.89                | - 16.84             | - 15.58                |
| $\Delta \varepsilon_i$             | 0.34          | -0.41                  | 0.37                | -0.16                  |
|                                    |               |                        |                     |                        |
|                                    | 5 00          | <b>⊿ 00</b>            | <b>⊿ Q</b> 1        | 4 87                   |

## HCOOH orbital energies (K=0.34)

|                                      | <b>10a'</b> | 2a''    | 9a'     | 1a''    | <b>8</b> a' | 7a'     |
|--------------------------------------|-------------|---------|---------|---------|-------------|---------|
| v <sub>rhola</sub> ;; (LDA)          | -27.33      | - 27.95 | - 27.16 | - 25.03 | - 28.02     | - 24.38 |
| $v_{xhole ii}$ (B)                   | - 2.63      | - 2.63  | -2.57   | -2.55   | -2.61       | -2.32   |
| v <sub>xresp,ii</sub> (GLLB)         | + 4.86      | + 5.64  | + 5.74  | + 5.86  | + 5.69      | + 5.64  |
| $v_{c,ii}$ (VWN)                     | - 1.83      | - 1.85  | - 1.83  | - 1.78  | -1.87       | - 1.78  |
| <sup>v</sup> xctot,ii                | - 26.94     | - 26.81 | - 25.82 | - 23.50 | - 26.79     | - 22.83 |
| $h_{ii} + v_{H,ii}$                  | +15.33      | +14.20  | +11.59  | + 8.06  | +10.20      | +5.68   |
| <b>E</b> <sub>i</sub>                | -11.61      | - 12.61 | - 14.23 | - 15.44 | - 16.59     | - 17.15 |
| <b>E</b> <sub>i</sub> (KS)           | - 11.51     | - 12.38 | - 14.48 | - 15.43 | - 16.64     | - 17.27 |
| $\Delta \varepsilon_i$               | -0.10       | -0.23   | 0.25    | -0.01   | 0.05        | 0.12    |
|                                      |             |         |         |         |             |         |
| <b>∆e</b> <sub>i</sub> (BVWN-<br>KS) | 4.23        | 3.82    | 4.07    | 3.78    | 4.17        | 3.82    |

## HCOOH orbital energies (K=0.34)

|                                    | 7a'     | <b>6a</b> ' | <b>5</b> a' | <b>4</b> a' |
|------------------------------------|---------|-------------|-------------|-------------|
|                                    |         |             |             |             |
| <i>v<sub>xhole,ii</sub></i> (LDA)  | - 24.38 | - 24.95     | - 31.36     | - 29.60     |
| $v_{xhole,ii}$ (B)                 | -2.32   | -2.24       | -2.08       | -2.05       |
| <i>v<sub>xresp,ii</sub></i> (GLLB) | + 5.64  | + 5.99      | + 5.91      | + 6.12      |
| $v_{c,ii}$ (VWN)                   | - 1.78  | - 1.79      | - 1.95      | - 1.90      |
| <sup>v</sup> xctot,ii              | - 22.83 | - 23.01     | - 29.48     | - 27.44     |
| $h_{ii} + v_{H,ii}$                | + 5.68  | + 2.17      | - 1.11      | - 5.41      |
| $\boldsymbol{\mathcal{E}}_i$       | - 17.15 | - 20.84     | - 30.95     | - 32.85     |
| <b>E</b> <sub>i</sub> (KS)         | - 17.27 | -21.17      | - 30.43     | - 32.94     |
| $\Delta \varepsilon_i$             | 0.12    | -0.33       | -0.16       | 0.09        |
|                                    |         |             |             |             |
| $\Delta \varepsilon_i$ (BVWN-KS)   | 3.82    | 3.88        | 4.14        | 4.14        |

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